

Deep Learning for Climate

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Climate and AI Workshop - Sorbonne Université (SU)

le cnam

Cédric



Outline

1 Context

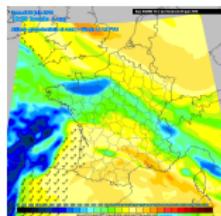
2 Neural Networks Models & Architectures

3 Deep Learning for Solar Irradiance Estimation

4 Perspectives

Big data in Climate and Beyond

- ▶ **Superabundance of data:** times series (sensor measurements), images (fisheye, satellite), spatio-temporal data (weather forecasts), videos, text, etc



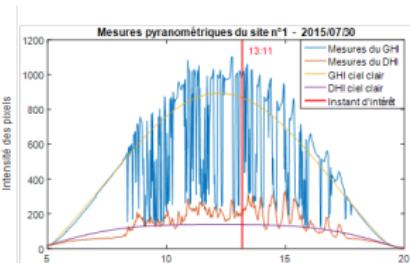
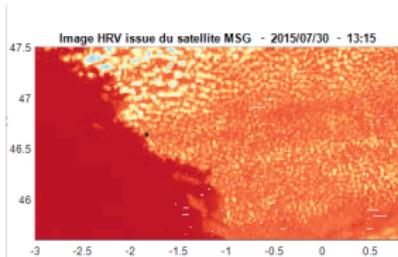
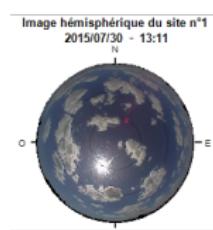
weather forecasts



Sensors, pyranometers



100M monitoring cameras

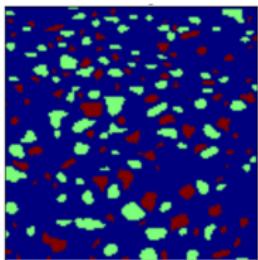
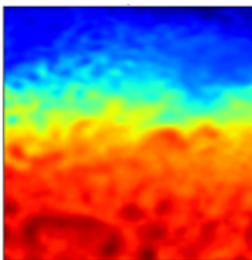
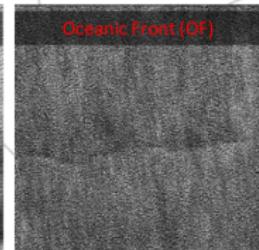
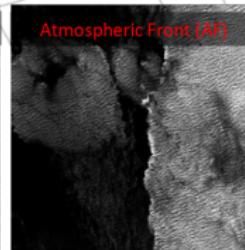
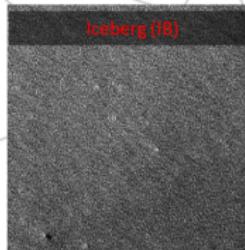
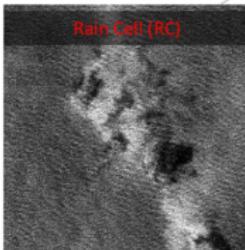
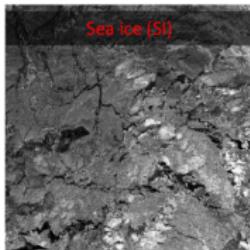


- ▶ Obvious need for **Artificial Intelligence** with these data
⇒ **Recognition, Decision Making**

Decision Making in Climate

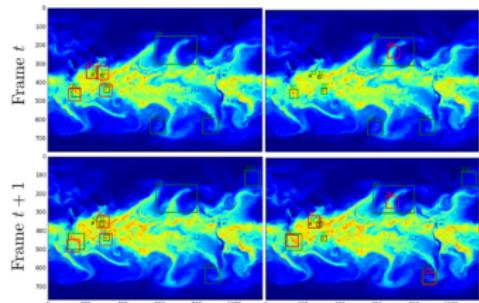
- ▶ **Huge number of applications:** classification, e.g. RADAR images, segmentation, e.g. eddies, forecasting, e.g. extreme weather event

[Chen et al., 2018b]



- Anticyclonic
- Cyclonic
- Non eddy

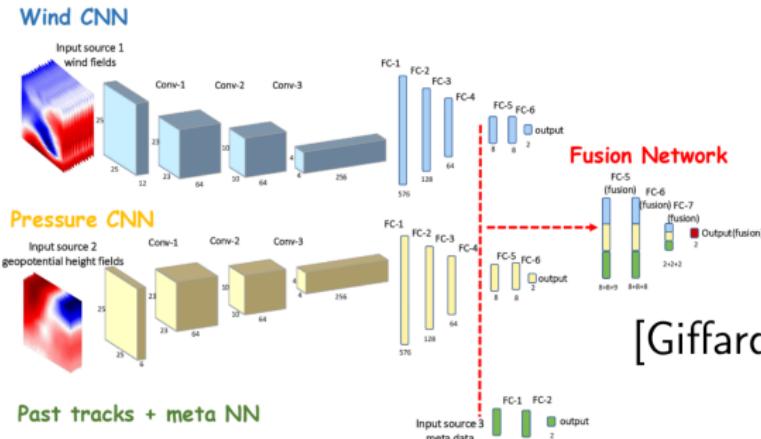
[Lguensat et al., 2018]



[Racah et al., 2017]

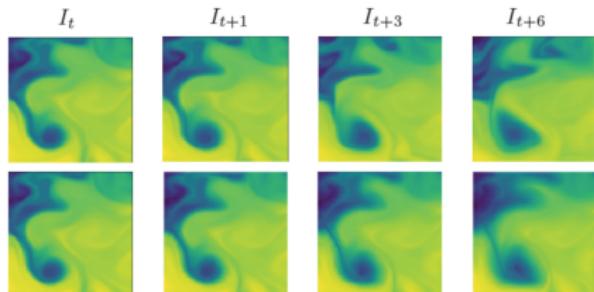
Decision Making in Climate

- Using Multiple inputs (wind, height, meta-data): hurricane track forecast



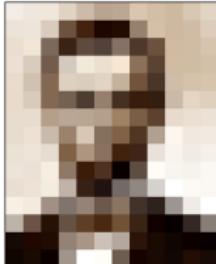
[Giffard-Roisin et al., 2018]

- Exploiting external knowledge: sea surface temperature prediction



[de Bezenac et al., 2018]

Recognition of low-level signals: filling the semantic gap



What we perceive vs
What a computer sees

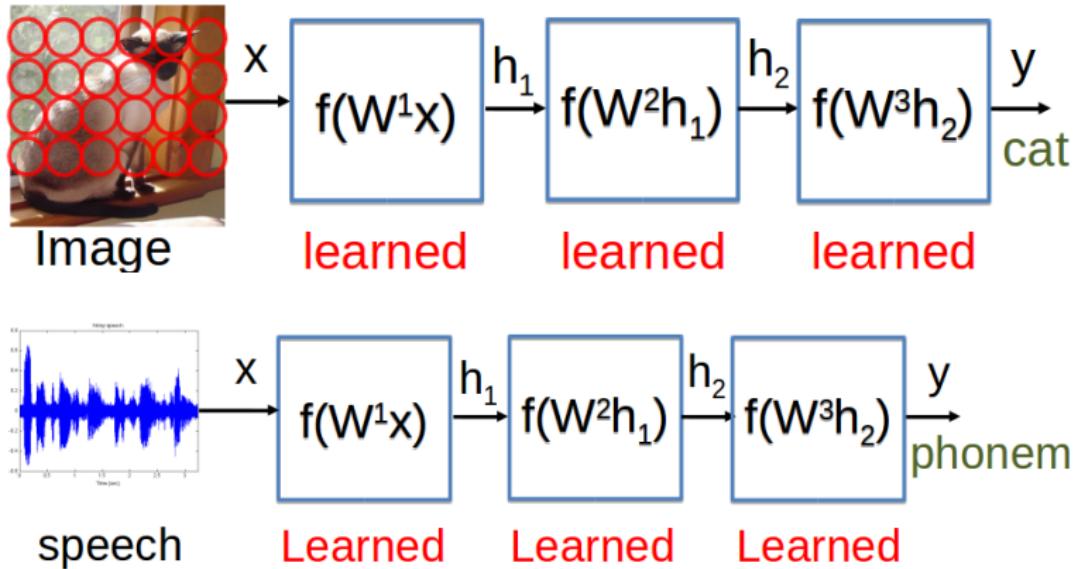


140	239	240	225	206	105	180	218	211	206	216	225
240	239	218	110	47	91	94	182	218	206	208	225
240	242	123	50	94	82	132	77	108	208	206	215
239	217	115	212	240	236	247	139	91	209	206	215
230	206	231	222	218	218	180	214	74	208	210	214
230	217	131	118	77	132	69	98	92	205	218	219
231	232	182	180	138	179	159	123	93	232	236	235
230	238	201	180	218	189	129	81	178	202	241	240
230	238	230	128	172	138	95	63	234	249	241	245
237	238	247	143	39	78	10	94	238	248	247	251
234	237	240	183	58	33	113	144	223	233	233	251
240	249	181	128	149	109	138	83	47	198	239	254
180	167	39	182	94	73	114	68	17	7	61	137
23	32	38	148	308	208	179	43	27	37	12	8
17	26	12	160	235	255	189	22	26	19	35	24

- ▶ Illumination variations
- ▶ View-point variations
- ▶ Deformable objects
- ▶ intra-class variance
- ▶ etc

⇒ Need for "good" Intermediate Representations

Deep Learning (DL) & Recognition of low-level signals



- ▶ **DL: learning intermediate representations**

- ▶ ⊕ **Deep**: hierarchy, gradual learning
- ▶ ⊕ Common learning methodology, few expert knowledge

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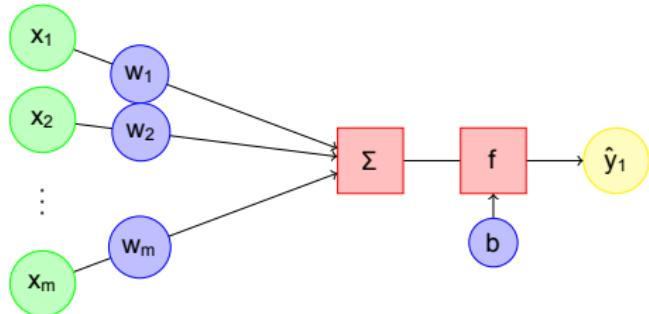
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Neural Networks (NN)

► The formal Neuron

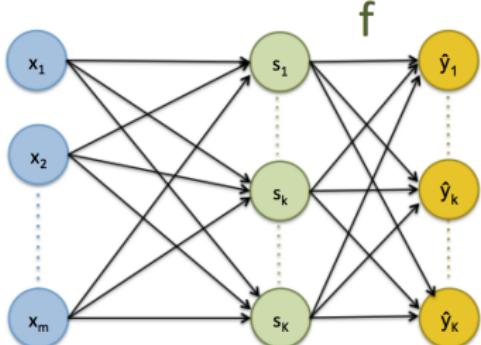


x_i : inputs
 w_i, b : weights
 f : activation function
 y : output of the neuron

$$y = f(w^T x + b)$$

Figure: The formal neuron – Credits: R. Herault

► Neural Networks: Stacking several formal neurons \Rightarrow Perceptron



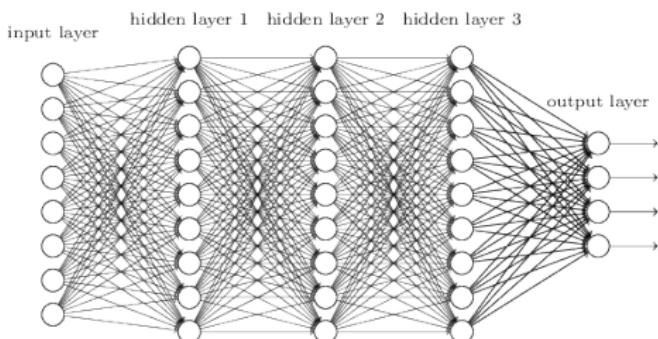
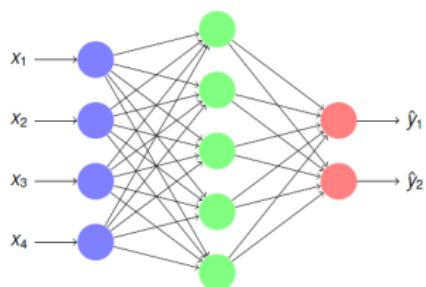
► Soft-max Activation:

$$\hat{y}_k = f(s_k) = \frac{e^{s_k}}{\sum_{k'=1}^K e^{s_{k'}}}$$

\Rightarrow Logistic Regression (LR) Model !

Deep Neural Networks (DNN)

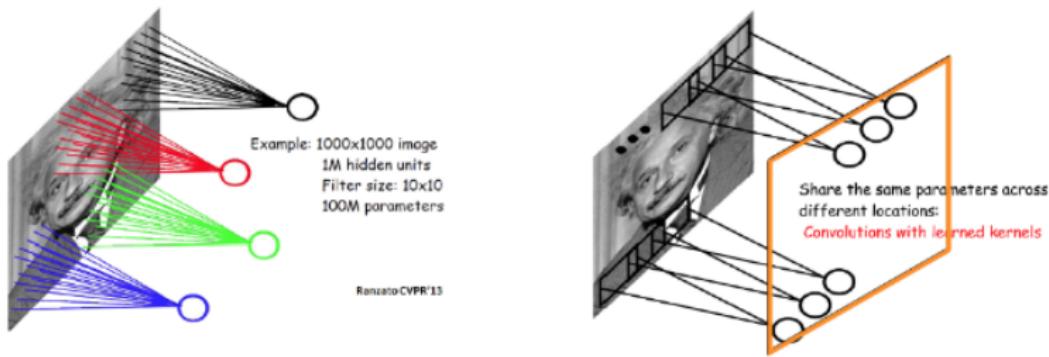
- ▶ **Multi-Layer Perceptron (MLP):** Stacking layers of neural networks
 - ▶ More complex and rich functions / Logistic Regression (LR)
 - ▶ **Neural network with one single hidden layer \Rightarrow universal approximator** [Cybenko, 1989]



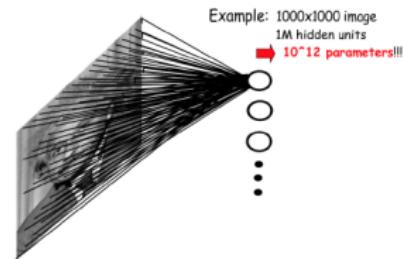
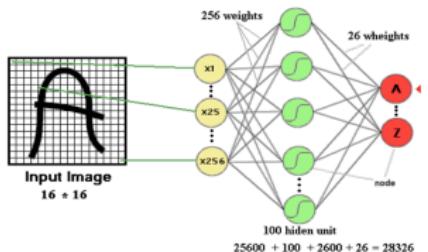
- ▶ **Basis of the "deep learning" field**
 - ▶ Hidden layers: intermediate representations from data
 - ▶ Can be learned with Backpropagation algorithm [Lecun, 1985, Rumelhart et al., 1986] (chain rule)

Convolutional Neural Networks (ConvNets)

- ▶ ConvNets: sparse connectivity + shared weights

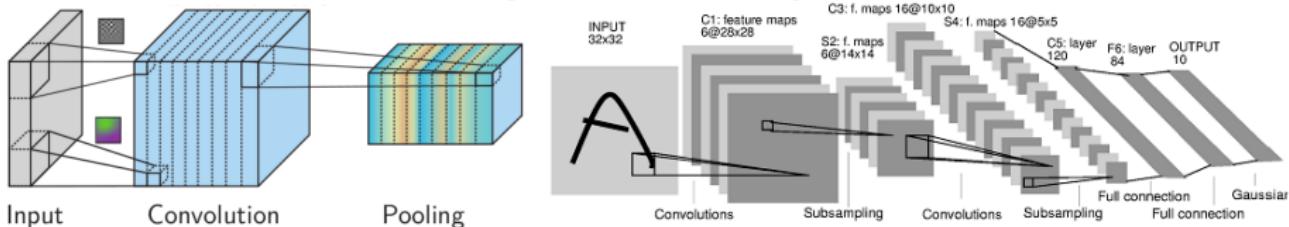


- ▶ Local feature extraction (\neq FCN)
- ▶ Overcome parameter explosion for FCN on images

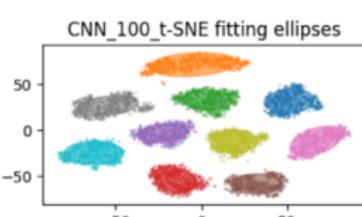
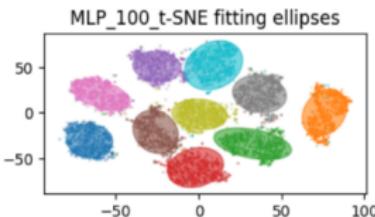
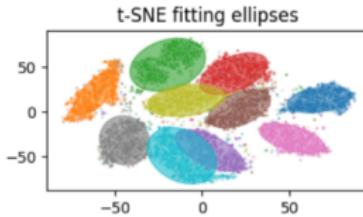
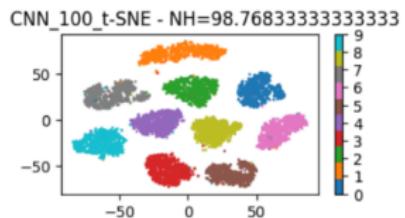
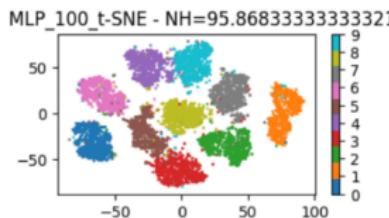
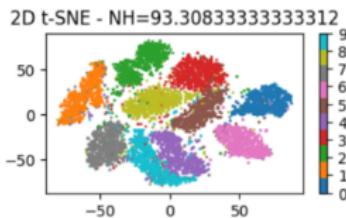


Convolutional Neural Networks (ConvNets)

- ▶ Elementary block: Convolution + Non linearity (e.g. ReLU) + pooling
 - ▶ Stacking: deep ConvNets [LeCun et al., 1989]

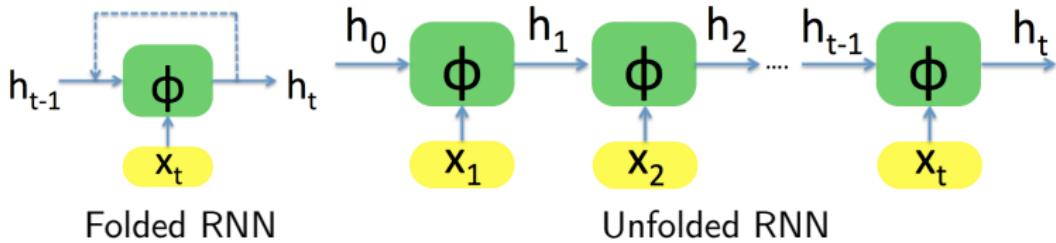


- Parameters ↓, invariance \Rightarrow ↑ generalization & manifold disentangling!

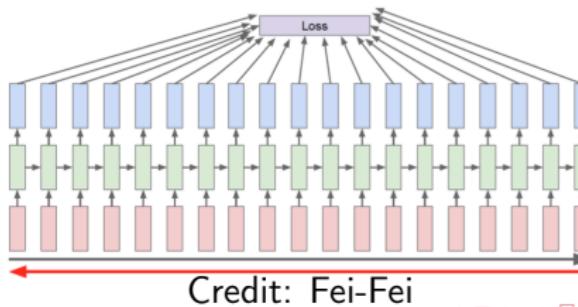


Recurrent Neural Networks (RNNs)

- ▶ **RNN Cell:** $\mathbf{h}_t = \phi(\mathbf{x}_t, \mathbf{h}_{t-1}) = f(\mathbf{Ux}_t + \mathbf{Wh}_{t-1} + \mathbf{b}_h)$ [Elman, 1990]
 - ▶ Loop, \mathbf{h}_t depends on current \mathbf{x}_t and previous state \mathbf{h}_{t-1}
 - ▶ \mathbf{h}_t : **network memory up to time t \Rightarrow Sequence processing**
- ▶ Universal program [Siegelmann and Sontag, 1995] approximators

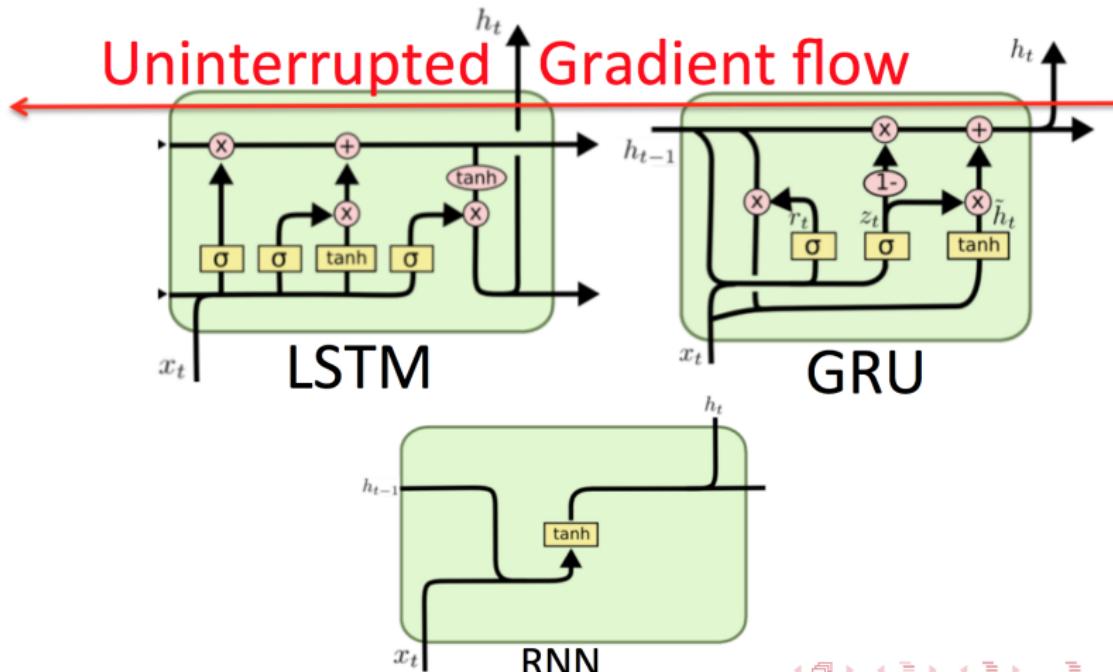


- ▶ Can be trained with Back-Propagation Through Time (BPTT)



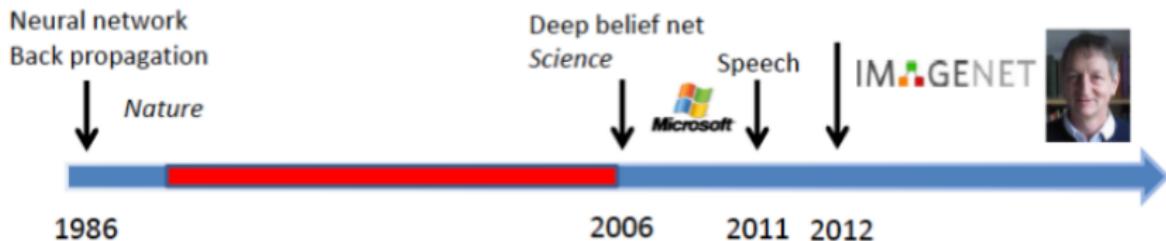
Recurrent Neural Networks (RNNs)

- ▶ BUT Back-Propagation Through Time \Rightarrow vanishing gradients
- ▶ Specific architectures:
LSTM [Hochreiter and Schmidhuber, 1997], GRU [Cho et al., 2014]
 - ▶ LSTM: Cell gate \Rightarrow uninterrupted gradient flow



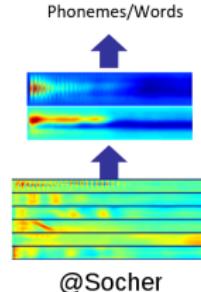
Deep Learning Success since 2010

- ▶ 90's / 2000's: difficult to train large ConvNets / RNNs on big data



- ▶ Deep Learning renewal since 2010
 - ▶ 2011: Speech Recognition

Acoustic model	Recog \\ WER	RT03S FSH	Hub5 SWB
Traditional features	1-pass -adapt	27.4	23.6
Deep Learning	1-pass -adapt	18.5 (-33%)	16.1 (-32%)



Deep Learning Success since 2010

- ▶ Deep Learning and ConvNet for Image Classification
 - ▶ ImageNet ILSVRC Challenge (Stanford):
 - ▶ 1,200,000 training images, 1,000 classes, mono-label
 - ▶ Based on WordNet hierarchy (ontology)
 - ▶ Up to 2012, leading approaches: handcrafted features + shallow ML (SVM)
 - ▶ **ILSVRC'12: the deep revolution**
⇒ outstanding success of ConvNets [Krizhevsky et al., 2012]

Rank	Name	Error rate	Description
1	U. Toronto	0.15315	Deep learning
2	U. Tokyo	0.26172	Hand-crafted
3	U. Oxford	0.26979	features and learning models.
4	Xerox/INRIA	0.27058	Bottleneck.

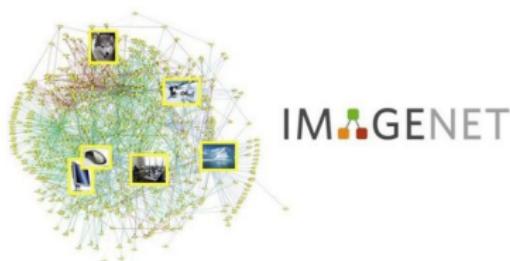
- ▶ RNNs SOTA for many sequential decision making tasks: speech, translation, text/music generation, times series, etc

2012: the deep revolution

Deep ConvNet success at ILSVRC'12

Two main practical reasons:

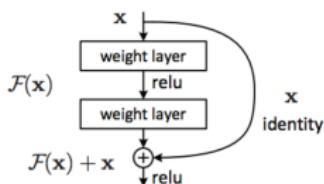
1. Huge number of labeled images (10^6 images)
 - Possible to train very large models without over-fitting
 - Larger models enables to learn rich (semantic) features hierarchies
2. GPU implementation for training
 - Relatively cheap and fast GPU
 - Training time reduced to 1-2 weeks (up to 50x speed up)



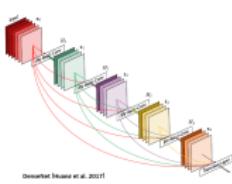
Current Trends in Deep Learning

Feature design \Rightarrow network architecture design

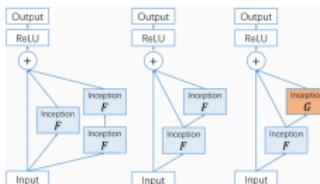
- Improved training properties, e.g. Res-Net or DenseNet \sim LSTM



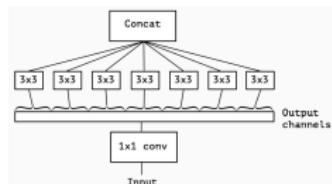
ResNet
[He et al., 2016]



DenseNet
[Huang et al., 2017]

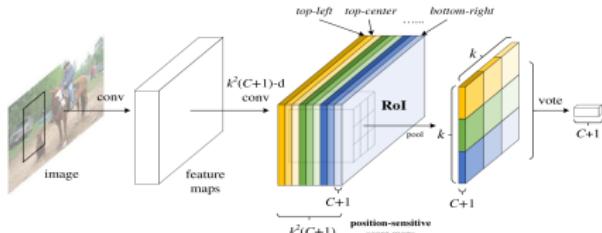


PolyNet
[Zhang et al., 2017]



Xception
[Chollet, 2017]

- Combining blocks for specific tasks, e.g. detection or ConvLSTM



[Dai et al., 2016]

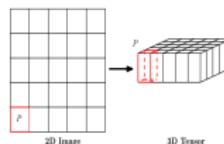


Figure 1: Transforming 2D image into 3D tensor

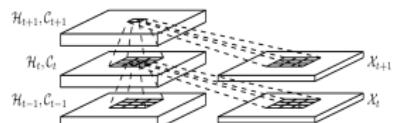


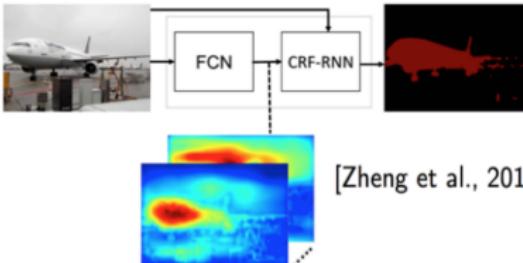
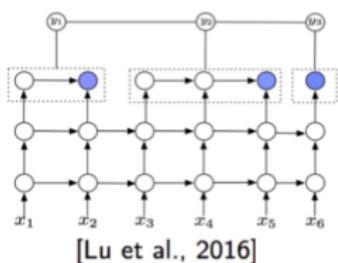
Figure 2: Inner structure of ConvLSTM

[Shi et al., 2015]

Current Trends in Deep Learning

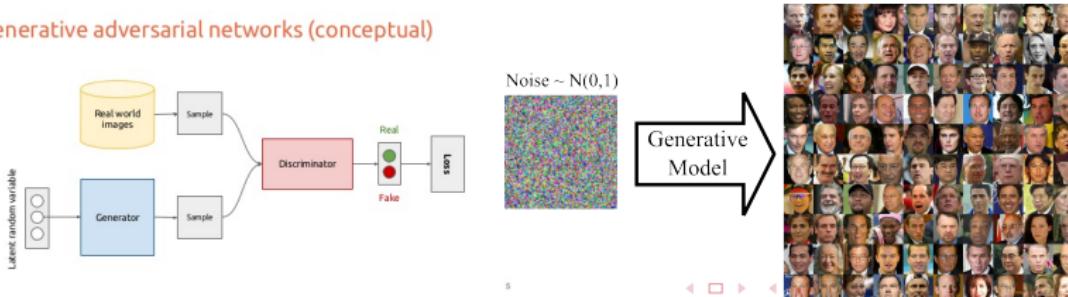
Training Models

- ▶ **Combining DL & structured prediction**, e.g. Conditional Random Fields (CRF)
 - ▶ Speech recognition (RNN+CRF), Semantic segmentation (ConvNets+CRF/RNN)



- ▶ **Generative Adversarial Networks**: Game Theory (generator vs discriminator)
 - ▶ Adversarial cost used beyond generation for distribution matching

Generative adversarial networks (conceptual)

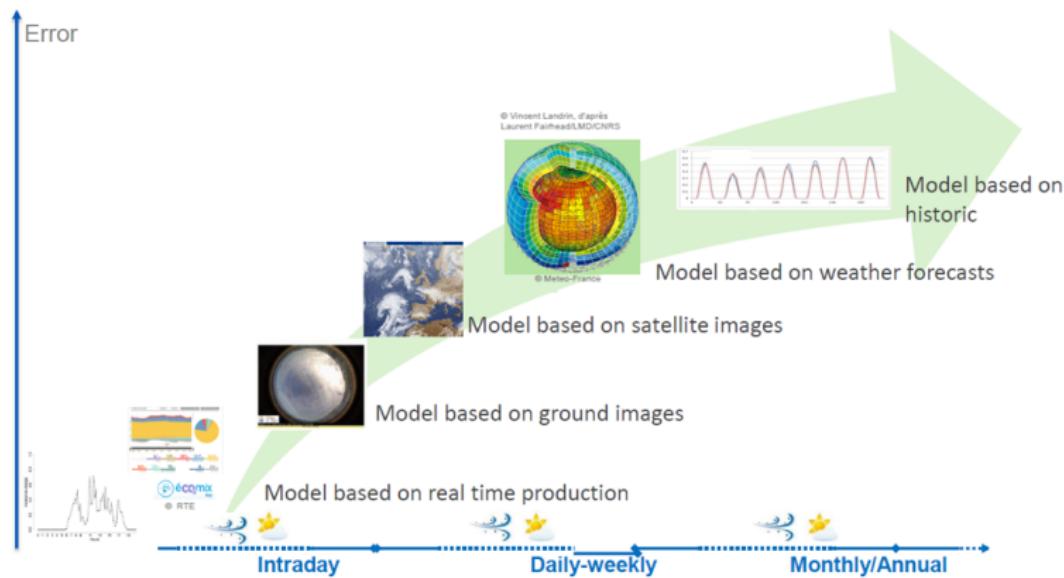


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Context: PhotoVoltaic (PV) energy forecasting

Different data sources for different horizons

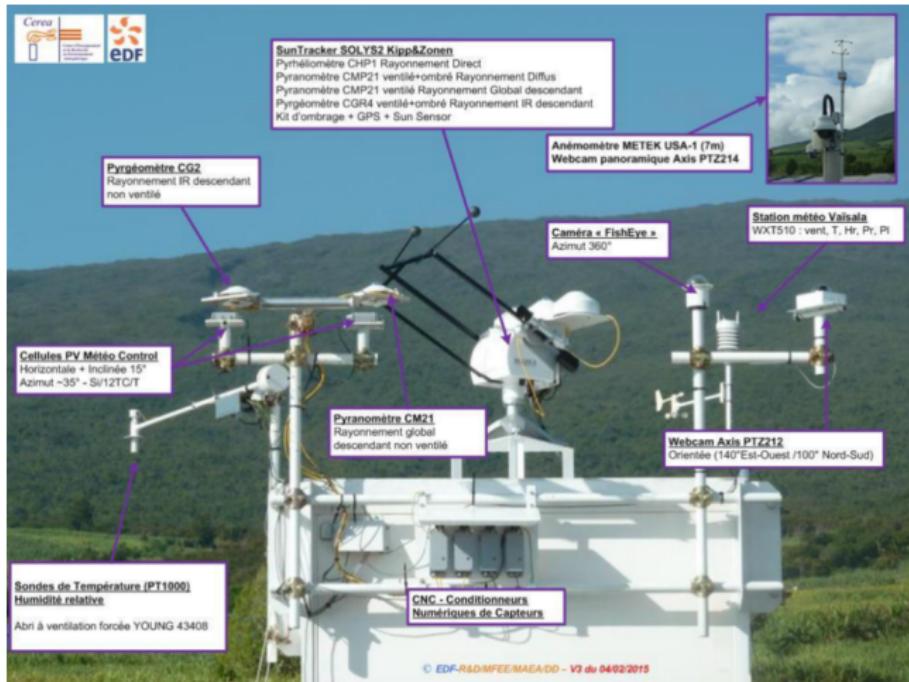


- ▶ **Ground based images:** very short term spatial & temporal horizons (0-20 min)
- ▶ **Application:** dynamic control of a hybrid system with PV, storage, diesel,...

Data

Meteorological campaign EDF R&D

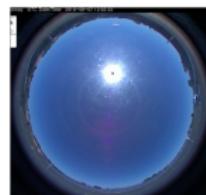
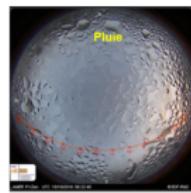
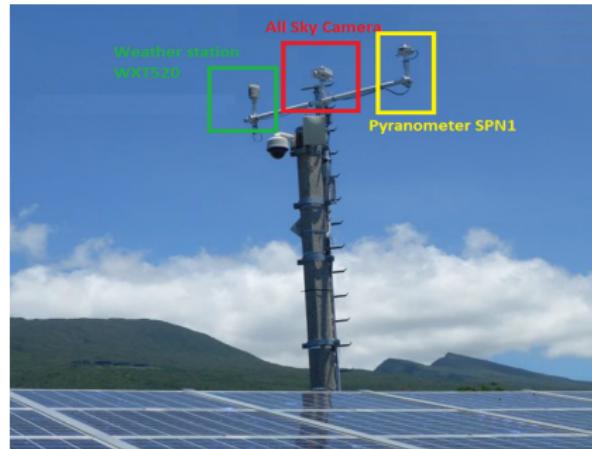
- EDF R&D experimental test site at La Reunion since 2012
- Many devices evaluated for solar resource assessment



Data

Meteorological campaign EDF R&D

- ▶ **Choice:** ground images + pyranometer
- ▶ **Goal:** Can we use low-cost cameras instead of pyranometers to estimate current and future solar irradiances?

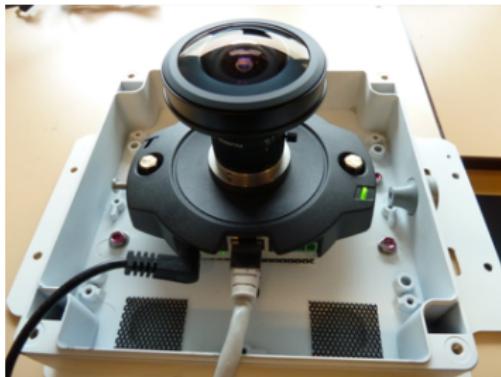


⇒ More than 7 Millions images and corresponding irradiation measurements collected since 2010

Data

Instrumentation

Fisheye camera: 180° hemispheric view of the sky, images every 10s

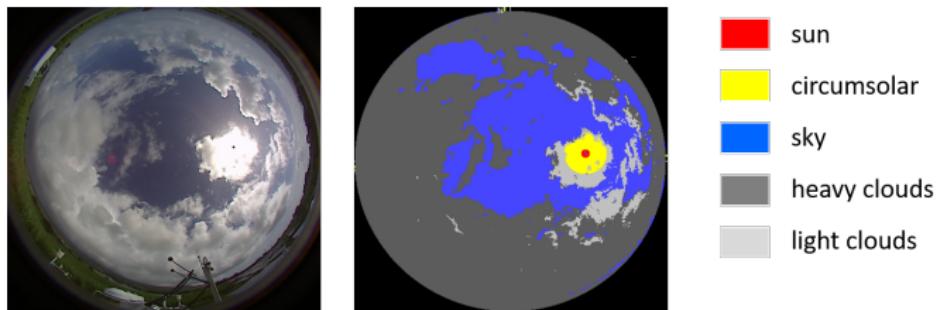


- ▶ **Pyranometer:** solar irradiance measurements every 10s
 - ▶ GHI: Global Horizontal Irradiance
 - ▶ DHI: Diffuse Horizontal Irradiance
 - ▶ DNI: Direct Normal Irradiance
- ▶ **Preprocessing:** irradiance values normalized by a clear sky model to remove seasonality ⇒ **KGHI**

Baseline

Irradiance estimation module

1. **Image segmentation:** with handcrafted thresholds on Luminance and R-B



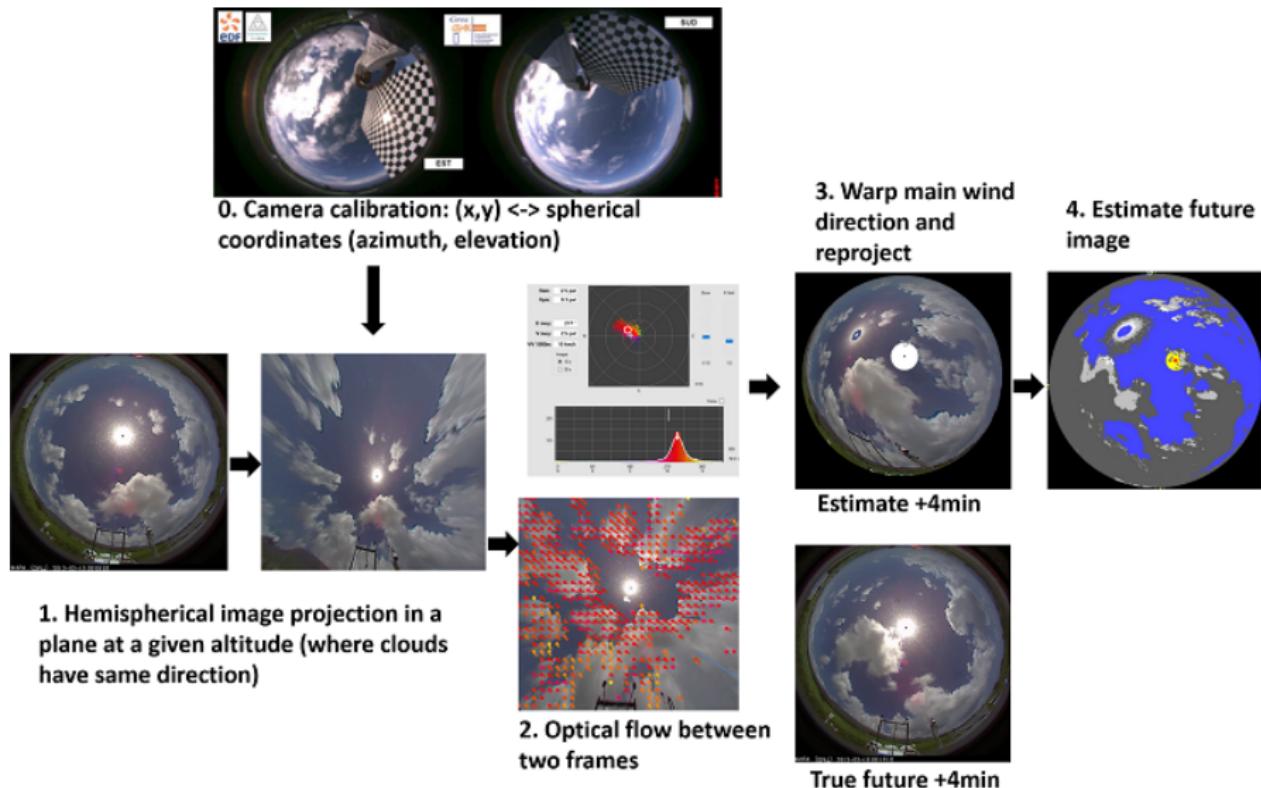
- ▶ Each image described with feature $x_i \in \mathbb{R}^5 \Leftrightarrow$ pixel class ratios
- ▶ **Database:** $(x_i, y_i)_{i=1:N}$ with x_i images and y_i corresponding KGHI

2. **Estimation:** kernel regression (Nadaraya-Watson model [Nadaraya, 1964]) for an unknown image x_0 :

$$\hat{y}(x_0) = \frac{C}{N} \sum_{i=1}^N e^{-\frac{\|x_0 - x_i\|^2}{2h^2}} y_i$$

Baseline

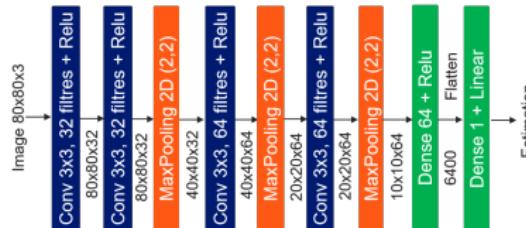
Forecasting module



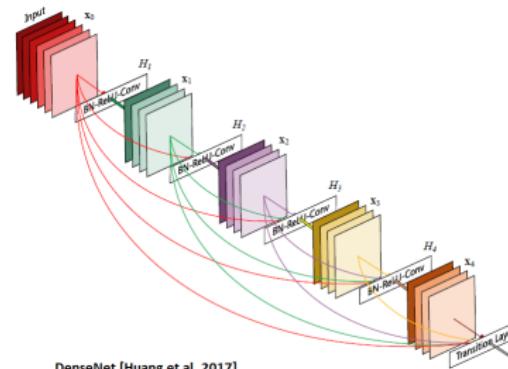
Estimating solar irradiance with ConvNets

Proposed neural network models

- ▶ **Small Convnet:** 475 000 parameters



- ▶ **Densenet model:** Densenet conv layers + dense regression layers
201 layers: 18 Millions parameters



DenseNet [Huang et al. 2017]

Estimating solar irradiance with ConvNets

Experiments

- ▶ **Experimental setup:**

- ▶ Training set: years 2012-2015 (4 190 064 images)
- ▶ Test set: year 2016 (1 265 717 images)

- ▶ **Implementation:**

- ▶ Python with Keras & Tensorflow backend
- ▶ Adam optimizer

- ▶ **Training time:** - Nvidia Quadro P6000 (24 Go RAM)

- ▶ 1 day for ConvNet, 6 days for DenseNet on a

- ▶ **Results for KGHI Estimation** (test set): large improvements of ConvNets wrt baseline

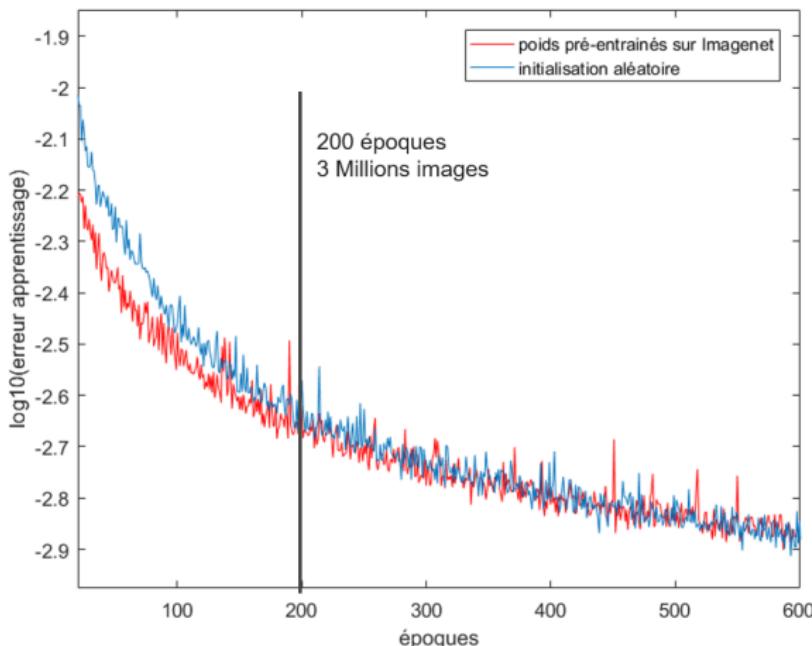
Model	Baseline	ConvNet	DenseNet
Mean Absolute Error (MAE)	0.1010	0.0448	0.0197
Normalized ¹ MAE	14.9 %	6.59 %	2.90 %
Root Mean Square Error (RMSE)	0.1467	0.06992	0.0328
Normalized RMSE	21.6 %	10.3 %	4.83 %

¹by the mean KGHI value over the training set

Estimating solar irradiance with ConvNets

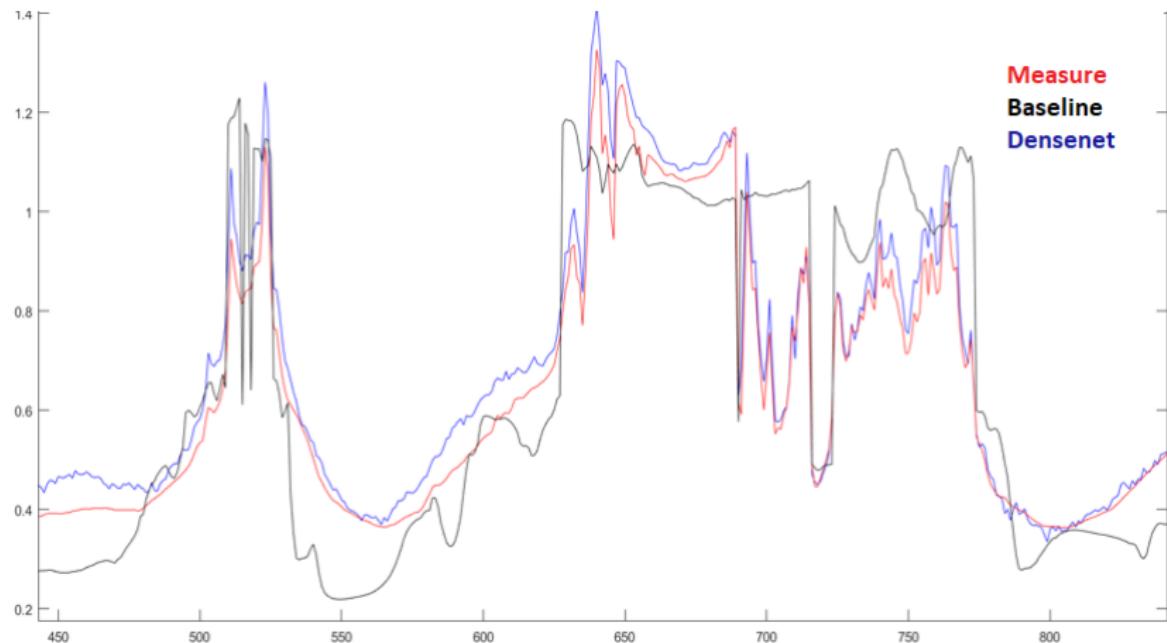
Training Evolution

Learning from scratch possible (ImageNet pre-training not necessary)



Estimating solar irradiance with ConvNets

Results on a particular day



Estimating solar irradiance with ConvNets

t-SNE visualization

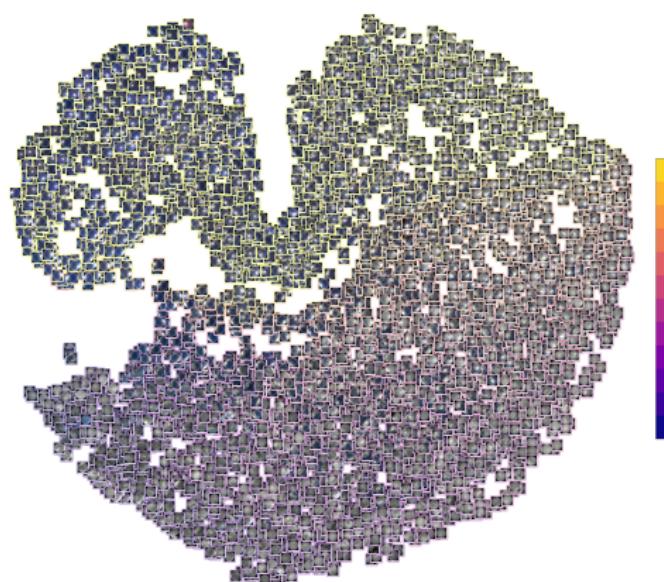
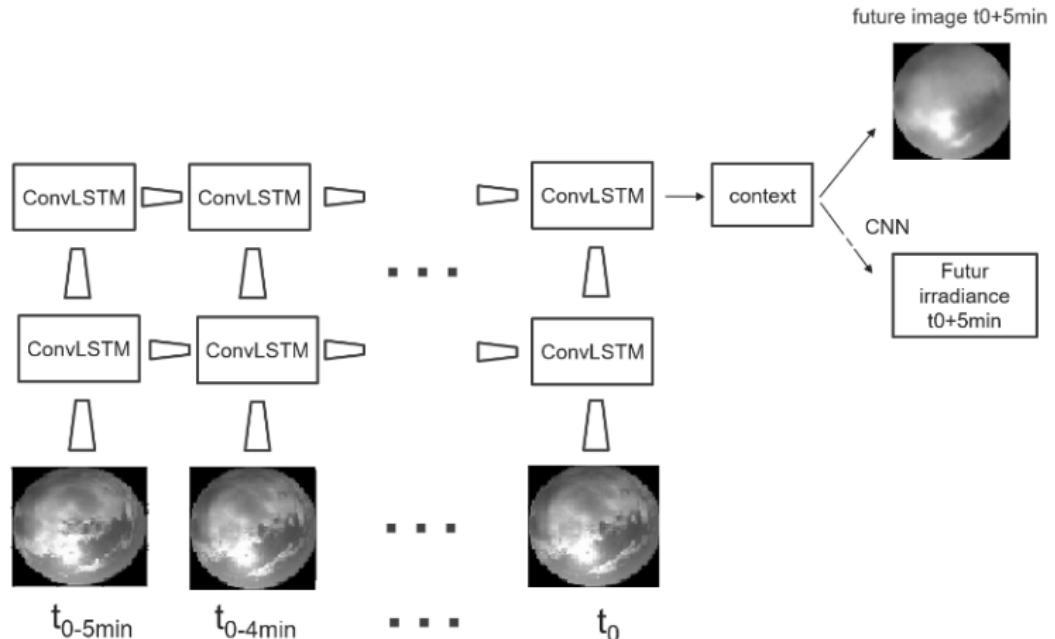


Figure: Clustering on Densenet features. Upper left: clear sky, upper right: cloudy, bottom: very cloudy, rainy

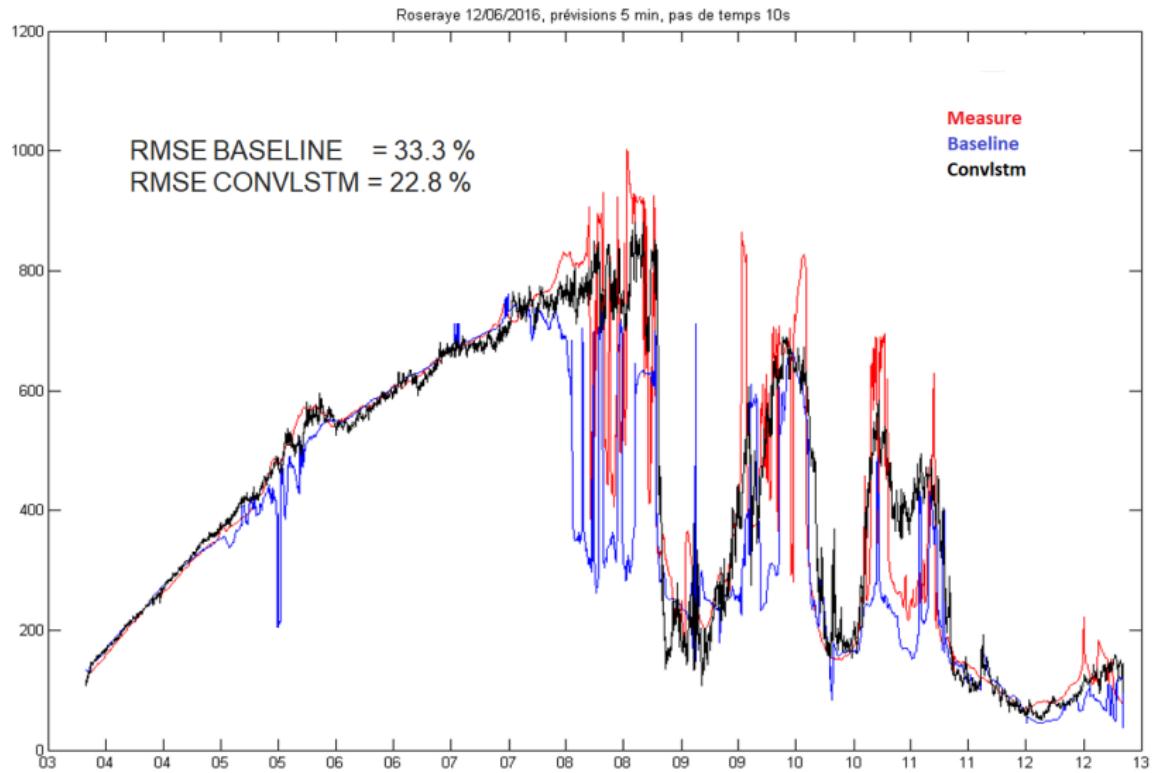
Ongoing work: irradiance forecasting

- ▶ **First proposed model:** light for computational purposes
- ▶ Input sequence: 10 grayscale 60x60 images every 30s
- ▶ Predict future image and irradiance at 5min
- ▶ Stacked ConvLSTM layers as spatiotemporal feature extractor



Ongoing work: irradiance forecasting

Preliminary results



Outline

1 Context

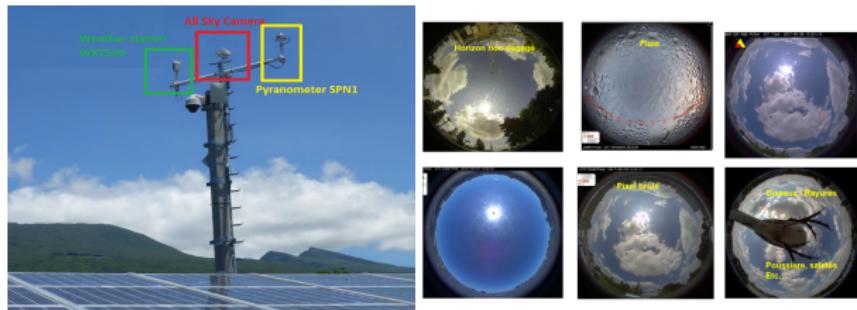
2 Neural Networks Models & Architectures

3 Deep Learning for Solar Irradiance Estimation

4 Perspectives

Conclusion & Perspectives

- ▶ Effective deep ConvNet solutions for irradiance predictions on static images
 - ▶ Favorable context: huge volume of annotated data



- ▶ Promising first results for future irradiance forecasting
 - ▶ Short-term: contribution of temporal information to irradiance & forecasts
 - ▶ Longer term: improve forecasting models and training methodologies

Forecasting future irradiances

Video Prediction

Deep learning for video prediction

- ▶ Direct RGB future image generation still challenging for large and complex natural images, predictions become blurry [Srivastava et al., 2015]
- ▶ To mitigate this: learn geometric transforms between images [Finn et al., 2016], use an adversarial loss instead of L2 [Mathieu et al., 2015]



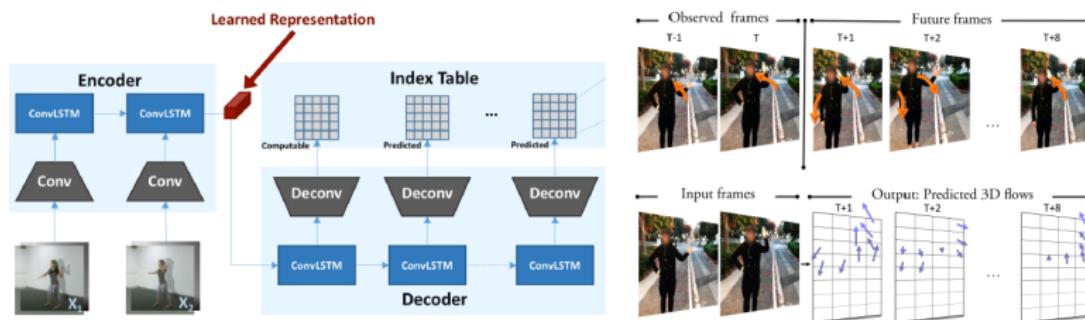
[Mathieu et al. 2015]

Forecasting future irradiances

Direct irradiance prediction

Direct irradiance prediction with deep learning without future image prediction

- Predict latent features in encoder-decoder RNN architectures [Luo et al., 2017]



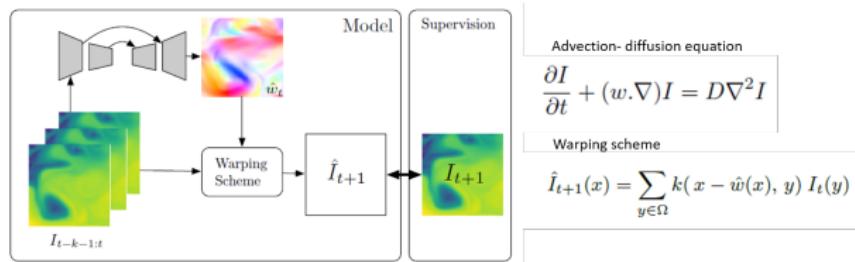
- Predict future image segmentations [Luc et al., 2018]

Forecasting future irradiances

Direct irradiance prediction

Prediction with physical knowledge

- Introduce a priori physical information (advection diffusion PDE)
[de Bezenac et al., 2018]



- Approximate differential equation solutions with neural nets
[Long et al., 2017, Chen et al., 2018a]

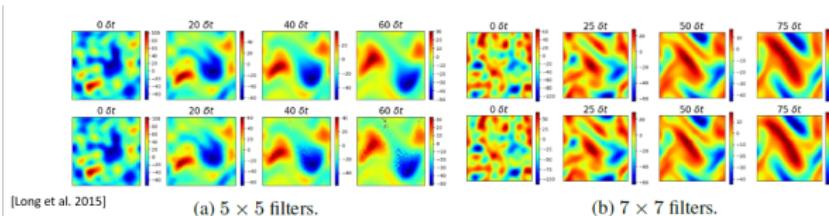


Figure 5: Images of the true dynamics and the predicted dynamics. The first row shows the images of the true dynamics. The second row shows the images of the predicted dynamics using the PDE-Net

Choice of loss function

How to penalize delays on the predicted irradiance time series ?

Goal: forecast irradiance ramps on time

Classical loss functions (MAE, RMSE) ill adapted to distinguish absolute value errors from temporal distortion errors. Possible solutions:

- ▶ signal gradient loss [Mathieu et al., 2015]
- ▶ loss based on Dynamic Time Warping [Cuturi and Blondel, 2017],...

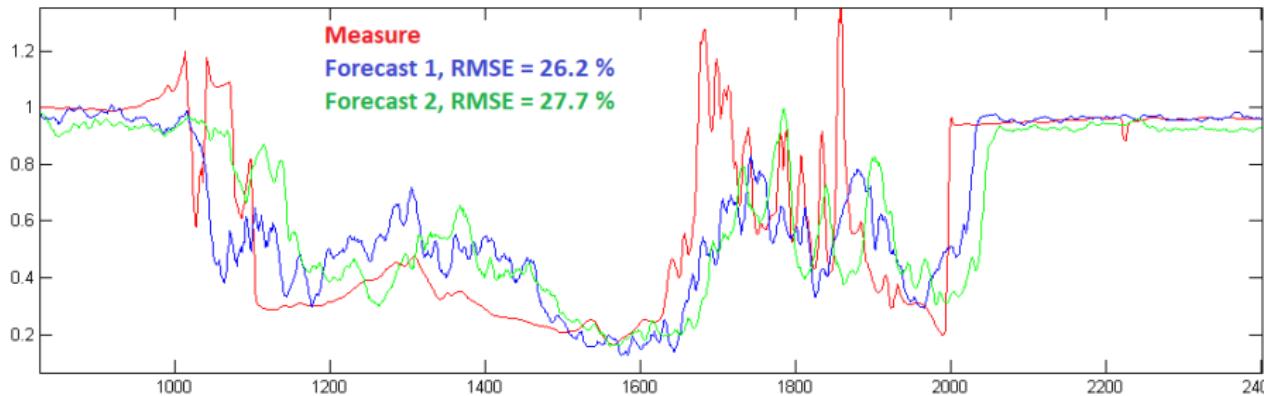
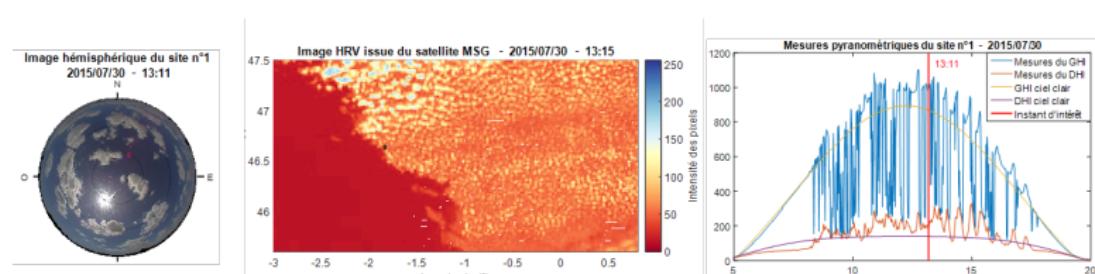


Figure: 2 forecasts with similar RMSE but different anticipating skills. Forecast 2 model seems always late

Multi-modal fusion

Ground images, satellite images, numerical weather forecasts

- ▶ Improve forecasting temporal and spatial scales by fusing different data sources:
 - ▶ PV production measurements
 - ▶ ground based camera
 - ▶ satellite image
 - ▶ numerical weather forecasts



- ▶ Use a network of multiple cheap cameras on a territory to anticipate global phenomena

Thank you for your attention!

Questions?

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